

*Some Considerations*

Of Mr. Nic. Mercator, concerning the Geometrick and direct Method of Signior Cassini for finding the Apogees, Excentricities, and Anomalies of the Planets; as that was printed in the Journal des Scavans of Septemb. 2. 1669: which Considerations are here deliver'd in the Latine Tongue, wherein they were written by the Author, as chiefly regarding the Learn'd in Astronomy, viz.

*Clarissimi Cassini Methodus*

*Investigandi Apogea, Excentricitates & Anomalias Planetarum, breviter Exposita & Demonstrata.*

Supponit Cl. Cassinus, ad Planetam in Ellipsi moventem extendi ab utroque foco duas rectas, quarum altera sit *medii*, altera autem *veri motus* linea. Constructio porro talis est;

<p><i>Fig. II.</i> L est Centrum Concentrici ABCDE. BLD est Diameter. BA, BC, BP, sunt intervalla apparentia. DE, DF, DQ, sunt intervalla medi- orum motuum. BE, BF, BQ; item DA, DC, DP, sunt lineæ rectæ. BE secat DA in H; BF secat DC in G; BQ secat DP in R.</p>	<p>RHG est linea recta. BI est perpendicularis ad RHG. I est Centrum Ellipseos. LI est Excentricitas. IO = LI. O est focus, circa quem ordinatur medius motus; L, circa quem verus. IM = IN = LB. M est Apogæon; N, Perigeon; BLM Anomalia vera.</p>
---	--

*Demonstratio.*

I. Illustrissimus ac Reverendiss. *Sethus Wardus*, quondam in Celeberr. Acad. *Oxon.* Professor Astronomiæ Savilianus, nunc Episcopus Sarisburiensis, in *Examine Astronomiæ Philolaica*, edito Oxon. A. 1653. c. 6. docuit Methodum, ex data Anomalia *media* Planetarum, investigandi *veram*; quæ est hujusmodi:

*Fig. III.* C, est Centrum Ellipseos AEP: F, focus, circa quem ordinatur medius motus. S, focus, circa quem ordinatur verus motus. A, Apogæon. P, Perigeon. E, Ero five Planeta. AFE, Anomalia media. ASE, Anomalia vera. FET, linea recta, ET = SE. ST est linea recta.

In  $\triangle$  SFT dantur, 1. SF distantia focorum: 2. FT = FE + ES = AP. 3. AFT, angulus externus, five Anomalia media, æqualis summae angulorum FST & T. Ergo inveniri potest FSE, five Anomalia vera, æqualis differentiæ Angulorum FST & T. Nimirum

Vt

Ut femi-summa laterum  $FT$  &  $FS$ , ad femi-differentiam eorundem ;  
Ita Tangens femi-summæ angulorum  $FST$  &  $T$ , ad Tangentem femi-differentiæ eorundem.

Sed femi-summa laterum  $FT$  &  $FS$  invenitur, substituendo pro  $FT$  æqualem  $AP$ , cujus semis est  $AC$ , qui additus  $CS$  semissi ipsius  $FS$ , facit Semis-summam  $AS$ , distantiam Planetæ maximam.

Tum, si ex femi-summa  $AS$  auferatur latus minus  $FS$ , restat femi-differentia laterum  $FA$ , æqualis  $PS$ , distantiæ Planetæ minimæ ; ut fit

*Regula ex Anomalia Media data inveniendi veram :*

Ut  $AS$ , distantia Planetæ maxima, ad  $PS$ , distantiam minimam ;  
Ita Tangens dimidiæ Anomaliæ mediæ, ad Tangentem dimidiæ Anomaliæ veræ.

*Corollar. I.* Si continuetur  $SE$  usque ad  $U$ , ita ut  $EU$  sit  $=$  ipsi  $FE$ , & tota  $SU =$  Axi  $AP$  ; erit  $\triangle FSU$  angulus  $U$  semis Prosthaphæreseos  $FES$ , ideoque æqualis femi-differentiæ angulorum Anomaliæ mediæ & veræ, h.e. ipsorum  $AFE$  &  $ASE$  ; & externus  $AU =$  femi-summæ eorundem  $AFE$  &  $ASE$  angulorum, ablata scil. femi-differentiâ  $UE$  ex majori  $AFE$ . Unde oriuntur duæ Analogiæ :

1. Ut Sinus femi-summæ Anomaliæ mediæ & veræ  $AU$ , ad Sinum femi-differentiæ eorundem,  $U$  ; Ita  $SU (=$  axi  $transverso AP)$  ad  $SF$ , distantiam focorum.

2. Ut Sinus femi-summæ Anomaliæ mediæ & veræ,  $AFV$ , ad Sinum Anomaliæ veræ  $FSU$  ; Ita  $SU$  (vel axis  $AP$ ) ad  $FU$ , subtensam Anomaliæ veræ : Ita quoque femi-axis  $AC$ , ad femi-subtensam  $UX$ , vel  $FX$ .

*Corollar. II.* Si in eodem Triangulo  $FSU$ , ex subtensæ  $FU$  puncto medio  $X$ , erigatur perpendicularis  $XE$  ; secabit illa  $SU$  in duas partes, quarum altera  $UE =$  est lineæ mediî motûs  $FE$ , altera verò  $SE$  est ipsa lineæ veri motûs.

*II. Fig. IV.* Sit  $a$  Centrum Concentrici  $chfi$ .  
 $cad$ , Diameter, eadêmque lineâ Apfidum.  
 $cb$ , Arcus Anomaliæ veræ, cui respondet  
 $di$ , Arcus Anomaliæ mediæ. Itaque

$cdh$ , est Angulus dimidiæ Anomaliæ veræ, &
$dci$ , Angulus dimidiæ Anomaliæ mediæ.
$ci$ & $dh$ sunt lineæ rectæ, secantes se mutuò in $g$ .

Ab Intersectionis puncto  $g$  demittatur ad  $cd$  perpendicularis  $gb$ . Erit igitur,

$db.bg ::$  Radius ad tang.  $bdg$  vel  $cdh$ .

Et  $cb.bg ::$  Rad. tang.  $bcg$  vel  $dci$ .

Ergo

Ergo  $db \times \text{tang. } cdb = bg \times \text{Rad.} = cb \times \text{tang. } dci$ .

Quare  $db \cdot cb :: \text{tang. } dci \cdot \text{tang. } cdb$ ; hoc est,  $db$  erit ad  $cb$ , ut tangens dimidiæ Anomalix mediæ ad tangentem dimidiæ Anomalix veræ; adeoque (per Regulam supra expositam) ut distantia Planetæ maxima, ad distantiam minimam. Quæ obrem  $db =$  erit distantia Planetæ maxima, &  $cb$ , minima, &  $ab$ , excentricitati.

Cùmque idem eodem modo demonstretur de cæteris omnibus Interfectionum punctis, nimir. Perpendiculares ab ipsis ad  $cd$  lineam incidere in punctum  $b$ ; oportet, ut recta, jungens ipsas Interfectiones, congruat perpendiculi  $bgf$ .

III. Ductâ diametro  $hak$ , fiat arcus  $kl =$  arcui  $id$ , & ducantur  $kc$  &  $hl$ , secantes se mutuò in  $p$ . Ab  $h$  in  $bgf$  demittatur perpendicularis  $hr$ , eadêmque parallela Apsidum lineæ  $ca$ ; erit angulus  $rhs$  semi-differentia arcuum Anomalix veræ  $ch$ , & mediæ  $di$ . Tum ab eodem  $h$  puncto ducatur recta  $hb$ , faciens cum  $kh$  angulum  $=$  angulo  $rhs$ , & occurrens lineæ Apsidum in  $\beta$ . Erit  $\Delta i a \beta h$  angulus  $\beta ah$  mensura arcûs  $ch$ , sive Anomalix veræ, &  $\beta ha$  semi-differentia Anomalix veræ & mediæ (ex Constructione;) & externus  $c \beta h$  (æqualis duobus internis & oppositis  $\beta ah$  &  $\beta ha$ , adeoque compositus ex Anomalia vera & semi-differentia ejus à media) erit semi-summa Anomalix veræ & mediæ. Ergo, per Corollarii I<sup>mi</sup> Analogiam priorem; Vt Sinus  $c \beta h$ , ad Sinum  $\beta ha$ ; ita Radius  $ah$ , ad Excentricitatem  $a^2$ . Sed supra demonstravimus quoque  $ab$  æqualem Excentricitati. Ergo punctum  $\beta$  congruit puncto  $b$ .

Tum ex  $b$  excitetur ipsi  $hb$  perpendicularis  $bt$ ; Aio, hanc continuatam incidere in punctum Interfectionis  $p$ . Nam Triangula  $rhs$  &  $bht$  sunt similia, ex Constructione; quemadmodum &  $\Delta^m hpk$  simile est  $\Delta o hgi$ , cum eidem peripheriæ  $cb$  insistentes anguli  $pkb$  &  $gih$  sint æquales, nec non æqualibus peripheriis  $kl$  &  $id$  insistentes anguli  $phk$  &  $ghi$  æquales; quare & tertius  $hpk$  æqualis est tertio  $hgi$ . Et ex æqualibus  $phk$  &  $ghi$  ablatis æqualibus  $bht$  &  $rhs$ , restant æquales  $phb$  &  $ghr$ . Vnde sic arguo:  $srh = tbb$ , &  $rhs = bht$ , Ergo  $hsr = htb$ ; ergo & Complementa horum ad semi-circulum sunt æqualia, nimir.  $rsi = btk$ ; &  $sig = tkp$ , Ergo &  $igs = kpt$ , quibus ablatis ex æqualibus  $igh$ , &  $kpb$ , restat  $hgs = hpt$ ; &  $ghr = phb$ , Ergo &  $hrh = hbp$ . Sed  $hrh$  est rectus, Ergo &  $hbp$  rectus est. Cum verò &  $hbt$  rectus sit ex Constructione, erit  $tb$  in directum ipsi  $bp$ . Cùmque idem eodem modo demonstretur de quavis alia Interfectione linearum ab  $h$  &  $k$  ad congruentiâ Anomalix veræ & mediæ puncta ductarum; patet, non modo rectam, jungentem interfectiones, transcurram per  $b$  punctum; sed &  $hb$ , lineam perpendicularem fore ad eandem Jungentem, q. e. dem.

*Corollarium.* Si à quovis puncto Anomalix veræ, puta  $b$ , ad respondens punctum Anomalix mediæ ducatur recta  $bi$ ; excitata è Centro Excentrici  $b$ , ipsi  $cbd$  perpendicularis  $bf$  secabit ipsam  $bi$  in  $s$  eâ ratione, quam linea mediæ motûs obtinet ad lineam veri motûs.

Nam per *Corollarium* I<sup>mi</sup> Analogiam posteriorem,  $hb$  est semi-subtensa; Ergò per *Corollarium* II<sup>um</sup>, perpendicularis erecta  $exb$ , nimir.  $bt$ , secat diametrum  $hk$  in  $t$  eâ ratione, quam linea mediæ motûs obtinet ad lineam veri motûs. Ergò &  $rs$  (sive  $bf$ ) secat  $hilineam$  eadem ratione in  $s$ ; propter demonstratam modò figurarum  $tbbkpb$  &  $srhighb$  similitudinem.

Cæterum ex laudata superius Reverendiss. *Wardi* Methodo inveniendi primam inæqualitatem, non est difficile, alium adhuc modum investigandi Apogæa & Excentricitates, non minus directum & Geometricum, & Observationes quovis admittentem, producere; quem & paucis exponam. Plures modos invenit Astrophili in Reverendiss. Viri *Astronomia Geometrica*, edita *A.* 1656, ad quam eos remitto. Interim

*Fig. V.* Sint  $l$  &  $d$  duo foci Ellipseos;  $t$  &  $u$  duo puncta veri motûs Planetæ; arcus Ellipseos  $tn$  ex  $l$  spectatus sub angulo  $tlu$ , & ex  $d$ , sub angulo  $tdu$ ; item distantia focorum  $ld$  ex  $t$  spectatus sub angulo  $dtl$ , & ex  $u$ , sub angulo  $dul$ : Aio, differentiam angulorum  $tlu$ ,  $tdu$ , a qualem esse differentiam angulorum  $dtl$  &  $dul$ .

Cùm enim trianguli  $lux$  tres anguli simul sumpti æquales sint trianguli  $dtx$  tribus angulis simul sumptis; si auferantur utrinque æquales  $lxu$  &  $dxr$ , reliquorum duorum summa  $ulx + lux$  erit = summæ reliquorum  $tdx + dtx$ , & ab his æqualibus summis si auferantur inæquales, v. g.  $ulx$  ex priori, &  $tdx$  ex posteriori; reliquorum,  $lux$  &  $dtx$ , differentia = est differentie ablatorum  $ulx$  &  $tdx$ ; quod erat propositum.

Centro  $l$ , intervallo axis transversæ  $mn$ , describatur Circulus  $abc$ , cujus arcus  $ab$  rursus ex  $l$  spectatur sub angulo  $alb$ , & ex  $d$ , sub angulo  $adb$ ; item distantia focorum  $ld$  ex  $a$  spectatur sub angulo  $lad$ , & ex  $b$ , sub angulo  $lbd$ . Ergò rursus differentia angulorum  $alb$  &  $adb$  = est differentie angulorum  $lad$  &  $lbd$ . Sed per *Coroll. I.* angulus  $lad$  semis est anguli  $lad$ , & angulus  $lbd$  semis anguli  $ltd$ . Ergò horum angulorum  $lad$  &  $lbd$  differentia = est semi-differentie angulorum  $lad$  &  $ltd$ ; ergò & angulorum  $alb$  &  $adb$  differentia = est semi-differentie angulorum  $ult$  &  $udt$ , quorum prior est intervallum apparens duarum Observationum, posterior autem, intervallum motûs mediæ. Datâ igitur horum intervallorum differentia, datur quoque hujus (*differentia*) semis, nimir. differentia angulorum  $alb$  &  $adb$ . Sed  $alb$  idem est cum  $ult$  dato; Ergò datur quoque  $adb$  angulus, sub quo periphæria  $ab$  spectatur ex  $d$ .

Simili modo ostendetur, differentiam angulorum  $tly$  &  $tdy$  æqualem esse summæ angulorum  $ltd$  &  $lyd$ ; nec non differentiam angulorum  $btc$  &  $bdc$  = esse summæ angulorum  $lbd$  &  $lcd$ . Cùmque  $lbd$  femis sit ipsius  $ltd$ , &  $lcd$  femis ipsius  $lyd$ ; erit sanè summa ipsorum  $lbd$  &  $lcd$  = semi-summæ angulorum  $ltd$  &  $lyd$ , hoc est, differentia angulorum  $btc$  &  $bdc$  = erit semi-differentiæ angulorum  $tly$  &  $tdy$ , quorum prior est intervallum apparsens duarum Observationum, posterior autem, intervallum motûs medii. Quare, datâ horum intervallorum differentiâ, datur quoque hujus femis, nimir. differentia angulorum  $btc$  &  $bdc$ . Sed  $btc$  idem est cum  $tly$  dato; Ergò datur quoque  $bdc$  angulus, sub quo periphæria  $bc$  spectatur ex  $d$ .

Unde liquet, ex datis intervallis Observationum mediis & apparentibus, dari angulos, sub quibus ex  $d$  spectantur Circuli  $abc$  periphæriæ quotvis, interceptæ à lineis veri motûs. Ergò, per *Herigoni Theor. Plan. l. 1. c. 3. Prop. 12. Schol. 1.* totidem Circuli segmenta describi possunt, capacia angulorum, sub quibus isti arcus conspiciuntur ex  $d$ , quæ segmenta omnia se mutuò interfecabunt in  $d$ . Possunt igitur & hac Methodo inveniri Apogæa & Excentricitates Planetarum, delineatione Geometricâ, adhibitis Observationibus quotvis; nec difficilius est, Circulos ducere, quàm lineas rectas.

Sed ut demus id, quod verum est, Clarissimi *Cassini* delineationem Geometricam non-nihil expeditiorem esse; verendum est interim, ne, si *επιβεβαιω* Astronomis expetitâ sectemur, Diagrammata requirat enormis magnitudinis, adeoque operosior evadat, quàm ipse Calculus. Ad hunc autem accedentes, utramque Methodum æquipollere deprehendemus.

Adhibeamus enim ex Observationibus *Tychonicis* tres, quæ *Dom. Cassini* Diagrammati quodammodo consentiant; nim. Observationem A, cùm *An. 1604, Mart. 28 d. 16 h. 23 m.* Mars observatus fuit in  $\approx 18 g. 37 m. 10 s.$  B, cum *An. 1587, Mart. 6 d. 7 h. 23 m.* idem Planeta visus fuit in  $\approx 20 g. 43 m. 0 s.$  Denique C, cùm *An. 1600 Jan. 18 d. 14 h. 2 m.* deprehenderetur in  $\approx 8 g. 38 m. 0 s.$  Est igitur inter A & B intervallum apparsens  $22 g. 54 m. 10 s.$  & huic respondens medium  $25 g. 58 m. 40 s.$ ; at inter B & C intervallum apparsens  $47 g. 5 m. 0 s.$  & medium  $56 g. 21 m. 57 s.$  Itaque

*Methodo*

( 1173 )

*Methodo Cassini, Fig. II.*

1. In Triangulo DBH,  
Dantur DB 10,00000  
DBH 12 | 99  
BDH 11 | 45  
Queritur BH 9,68106

2. In Triangulo DBG.  
Dantur DB 10,00000  
DBG 28 | 18  
BDG 23 | 54  
Quer. BG 9,70653

3. In Triangulo HBG.  
Dantur BH 9,68106  
BG 9,70653  
HBG 41 | 17  
Quer. BGH 64 | 95  
Cujus Compl. GBI 25 | 05  
Si auferas ex GBD 28 | 18  
Restat IBD vel IBL 3 | 13

4. In Triangulo GB I.  
Dantur BG 9,70653  
GIB 90  
GBI 25 | 05  
Quer. BI 9,66363

5. In Triangulo IBL.  
Dantur BI 9,66363  
BL (semis  $\tau$  BD) 9,69897  
IBL 3 | 13  
Quer. BLI 32 | 31, An. vera, & LI, 8,67284, Excentricitas.

*Methodo Herigoni, Fig. V.*

1. In Triangulo dbh,  
Dantur db 10,00000  
adb externus 24 | 44  
bhd 11 | 45  
Quer. bh 10,31894

2. In Triangulo dbg,  
Dantur db 10,00000  
cdb externus 51 | 72  
bgd 23 | 54  
Quer. bg 10,29347

3. In Triangulo hbg,  
Dantur bh 10,31894  
bg 10,29347  
hbg 41 | 17  
Quer. hbg (vel hbi) 64 | 95 = bsg  
Et hbi = sgb = 90:  
Ergo hbi = gbs = 25 | 05  
Ex gbi = gbs + sbi (= hbg - hbi) = 16 | 12  
Aufer dbh = hbi - dbi = 12 | 99  
Restat gbs + sbi - hbi + dbi = sbd (vel dbl) 3 | 13

4. In Triangulo gbs  
Dantur bg 10,29347  
bgs 90  
gbs 25 | 05  
Quer. bs 10,33637

5. In Triangulo dbl,  
Dantur bd 10,00000  
bl (semis  $\tau$  bs) 10,03534  
dbl 3 | 13  
Querit. bld 32 | 31 Anom. vera  
Et ld 9,00926 Excentricitas.

Nimir. Ut Fig. II. BL 9,69897, ad LI, 8,67284;  
Ita Fig. V. bl 10,03534, ad ld 9,00926.

Ex loco apparenti secundæ Observationis  
auferatur angulus Anomalix veræ B L I  
Restat locus Apogei

s.	g.	m.	sec.
5	25	43	0
1	2	18	36
4	23	24	24

Erat autem reverà ævo *Tychoñis* Apogeon *Martis* in  $\Omega$   $28\frac{1}{2}$  d., à quo deficit iste locus, calculo inventus, solidis quinque gradibus. Porrò, Ut B L 9, 69897, } Ita 5, 18290 Log-us 152369 distantia med.  $\delta$  tis, ad L I 8, 67284; } ad 4, 15677 Log-um 14347 Excentricitatis  $\delta$  tis.

Est autem vera Excentric.  $\delta$  tis 14179, quam ista, calculo inventa, excedit  $\frac{168}{14179}$  particulis.

Cæterum in ratiocinio secundum utramque Methodum instituto notare licet non modò perpetuam Triangulorum similitudinem, sed & Epilogismi congruentiam; ne quis Apogei & Excentricitatis sic inventæ à vero discrepantiam censeat errori Calculi imputandam. Sed nec Observationum vitio contingit; quas in dubium vocare nil aliud foret, quàm principia in Astronomia negare. Itaque restat, ut Hypothesin excutiamus.

Et *Ellipticæ* quidem Orbitæ Inventio sine controversia *Keplero* debetur; sed quibus Accelerationis & Retardationis gradibus incedant Planetæ, definire, non minùs pertinet ad integrandam Hypothesin, quàm ipsius Orbitæ determinatio. Quanquam autem ex Cl. *Cassini* (vel Interpretis ejus) sermone id nusquam apparet; attamen ex Constructione Problematis, & ejus Analyfi, manifestum est, eum supponere, Planetam ex foco superiori videri prorsus æquabili motu incedere. Fuit sanè, cum idem existimaret *Keplerm*, quod ejus Scripta evolventibus liquere potest. Sed cum id Observationibus nequaquam congruere animadverteret, mutavit sententiam, & lineam veri motùs Planetæ æqualibus temporibus æquales areas Ellipticas verrere professus est: Punctum autem, ex quo Planeta exactè æquabili motu procedere videtur, nullum omnino extare in hoc Universo, nisi id libratile statuere libeat. Nulli interim puncto propriùs æquabilem videri incessum Planetæ, quàm ipsi foco superiori Ellipseos. Neque inventus fuit hæctenus, qui areas *Kepleri* phænomenis satisfacere posse negaret; sed, cum eas Calculo directo exhibere nec ipse nec post eum quisquam potuerit, causati sunt nonnulli, *Keplerum*, nimis indulgentem causis *Physicis*, à *Geometria* diversum abiisse; quasi causæ physicæ repugnent *Geometricæ*, aut minus *Geometricum* sit Problema, quod, nullâ injectâ physicarum causarum mentione, sic proponitur: *Data area Trilinei, inter lineas absidum, & veri motus, nec non peripheriam Ellipticam intercepti, invenire Angulum ad Solem.* Habent igitur à *Keplero* responsum, qui illi ἀγεωστρίαν objiciunt; nim. *Eant ipsi & Schema solvant.*

Quamvis autem religio fuerit *Keplero*, ab Hypothesi, quam *Naturalem* esse planè persuasum habebat, recedere; quidni liberum foret aliis periculum facere, num via quævis alia detur, inæqualitatem Planetarum primam directo Calculo investigandi? Ideoque Vir Clariss. *Ism. Bullialdus* aggressus est ratiocinio *Geometrico* indagare, quâ semitâ, & quibus intentionis ac remissionis gradibus conveniret Planetas ferri, ut ab æquabili incessus norma, Astronomis ante *Keplerum* assumptâ, ad eam, quam spectamus, Inæqualitatem perduceremur. Perennant Illustrissimi viri  
monu-

monumenta , unde omnem hujus Inventi rationem haurire licet Astrophiliis. Amplexus eandem Reverendis. *Seb. Wardus*, primum ostendit, paria facere cum linea æquabilis motus circa alterum Ellipseos umbilicum gyrata; deinde & Calculi directi methodo ornavit eam, quam paulò antè recitavimus: Ita ut nil amplius desiderari posset, quàm ut *Urania* felicibus cæptis annueret. Cujus quidem nomine suscipere ausus fuit Illustris. Comes *Paganus*, edito, *biennio post*, ejusdem ferè tenoris Scripto, adeò veram esse Hypothesin, ut deprehensam circa Octantes discrepantiam, Astronomorum insectiæ tributam mallet. At Cl. *Bullialdus*, audiendam potius ipsam Astronomiam ratus, Observatorum ore loquentem, secundis curis, adhibita prioribus Inventis limitatione quadam, discrepantiam illam exterminavit. Unde porrò intelligitur, Hypothesin illam, cui Cl. *Cassinus* investigationem Apogeorum & Excentricitatum superstruit, tantundem ferè deficere à vero, quantum Cl. *Bullialdi* limitatio poterat, atque ab illo defectu pullulare eum quem supra notavimus, Calculi à Cælo dissensum.

Tantum vero abest, ut de Eximii Viri Inventione vel minimum delibatum velim, ut quicquid hujus lucubratiunculæ non hausi ex Reverendiss. *Wardo*, vel *Herigono*, id omne Ipsi libentissimè acceptum referam, qui ansum nobis præbuit hæc altius considerandi. Nec dubitamus, quin omnia ista multò uberius ac luculentius in promisso *Tractatu* exposita propediem reperturi simus, cujus Editionem maturam, pro eo quo flagramus divinissimæ Scientiæ amore, perquam avidè exspectamus.

An Account of Three Books.

I. *Esperienze intorno alla Generazione Degli' Insetti, fatte da Francisco Redi, Accademico della Crusca. In Firenze, A. 1668. in 4o.*

**T**HE Learned and Ingenious Author of this Book, lately come to the Publishers hands, though not yet (which is much disliked by the curious) into our Stationers Shops, doth with much industry undertake therein to evince, that there is no such thing as *Æquivocal Generation* but that every Animal is generated by the seed of another Animal, (its parent,) or, at least, from some Living and un-corrupted Plant, as out of Oak-Apples, and several Protuberances and Excrecencies of Vegetables.

*First* then, in the asserting of the *Universal* and true Generation of Insects by a peculiar and paternal Seed, the Author positively affirms, that he could never find, by all the Experiments and Observations, he ever made (of which he relateth a great number, by himself made upon all sorts of Animals) that ever any Insects were bred from Flesh, or Fish, or *putrified* Plants, or any other Bodies, but such, as Flies had access unto, and scatter'd their seed upon; he having taken extraordinary care and pains to observe, that alwayes on the Flesh, before it did verminate, there late Flies of the self same kind with those, that were afterwards produc'd thence; and again, that no Worms would ever come from any Flesh in Vessels well cover'd, and defended from the access of Flies; so that to him there is no generation of Insects from any dead Animals, but such as have been fly-blown.

And least it should be objected, that the reason, why in vessels exactly clos'd, no Insect breeds, is the want of Air, necessary to all Generation, He hath carefully covered several vessels with very fine Naples-vaile, for the Air to enter, though Flies could not; but that no worms at all were bred there, notwithstanding that many Flies swarmed about them, invited by the smell of the Flesh inclosed therein.

*Secondly*, to make out the other part of his Position. *viz.* That these Animals that are not bred by the seed of other Animals, are produced from some live Plant, or its

Excre-



Fig. II.

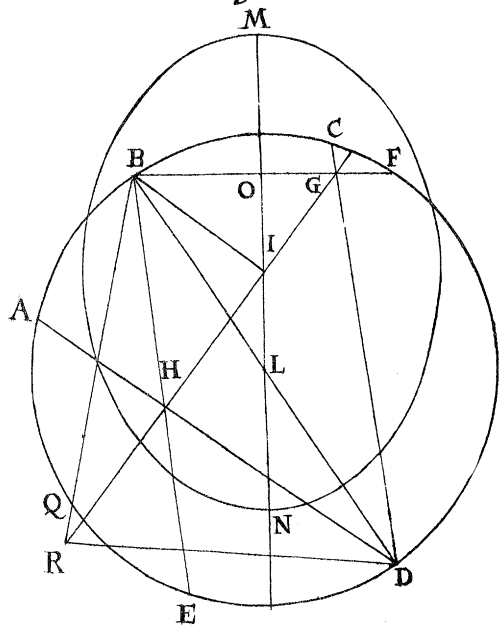


Fig. III.

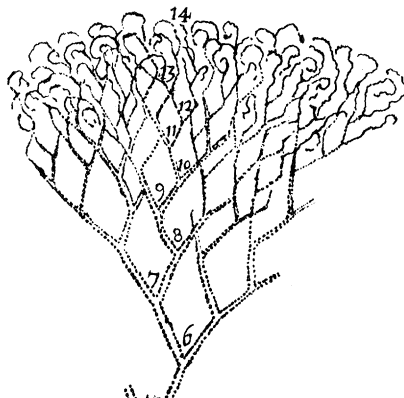
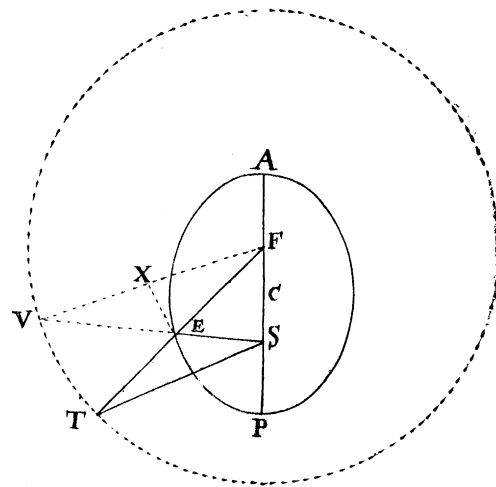


Fig. I.

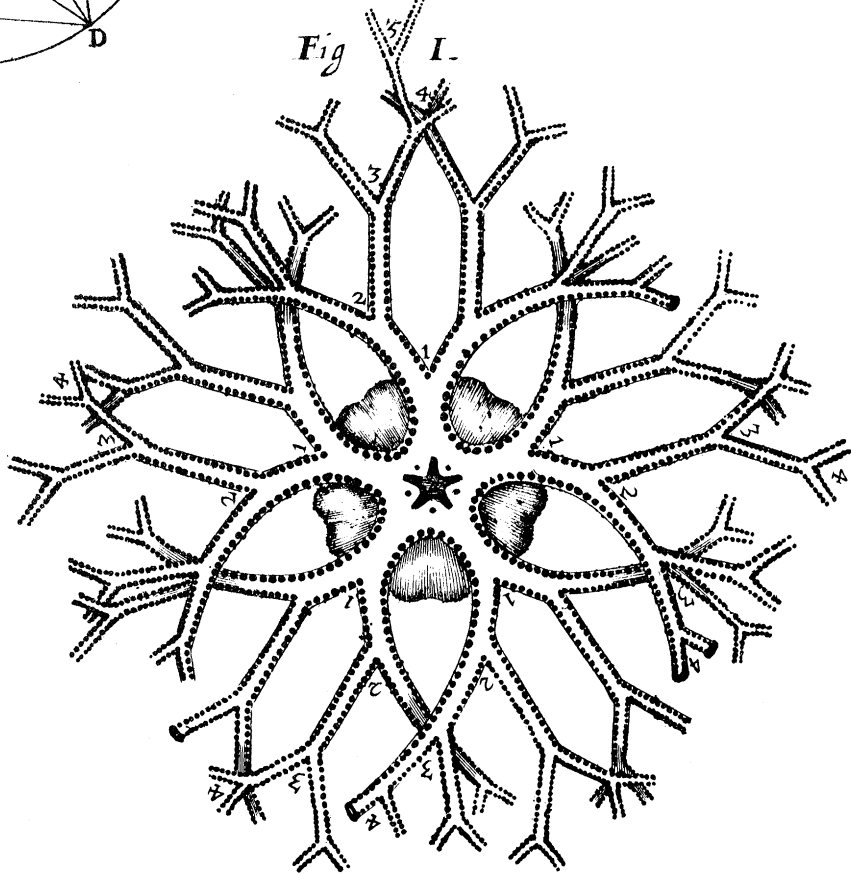


Fig. IV.

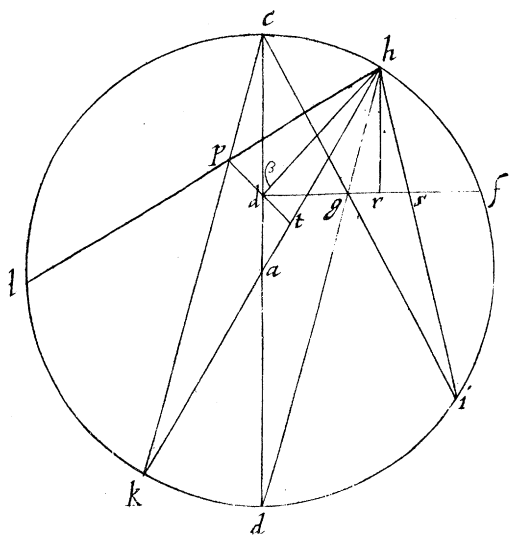


Fig. V.

